

AERONAUTICAL EPHEMERIDES

(Lecture delivered by Ingénieur Hydrographe Général COT, Hydrographer of the French Navy, before the IVth International Hydrographic Conference, 19th April, 1937).

The problem of fixing the position in aerial navigation is fraught with grave obstacles. In general, aviators are satisfied to observe the earth beneath them and to identify the points over which they are flying and which are shown on the charts at their disposal :— localities, rivers, hills or other irregularities of the terrain. This procedure is very convenient when it is possible to see and mark on the chart the details to be identified. This, however, is not always the case, for instance over the ocean or above the desert, or even in foggy weather or at night, and particularly in the latter case when lights are extinguished.

For this reason many naval officers who serve as pilots or observers in the French Air Service have sought to apply to aerial navigation the maritime procedure, i. e. astronomic method, which consists, as we know, in measuring the height of a celestial body above the sea horizon and noting the time at the same instant. Thereupon one computes the estimated altitude corresponding to the assumed latitude and assumed longitude; the difference between this and the observed altitude (taking into consideration the various sources of error : refraction, parallax, etc.) permits the line of position to be plotted.

In air-planes, however, the observers encounter rather special conditions. At high altitudes their hands become so chilled that they are obliged to wear thick gloves. Not only are they inconvenienced by the cold but also by the wind, as they are exposed to violent air currents, especially if flying at speeds of 150, 200 or 300 kilometres an hour (90, 140 or 210 miles per hour). Thus aviators demand that the calculations should be reduced to a minimum. If it becomes necessary to perform even a simple addition by writing $A + B$, that alone is regarded as considerable. All computations must therefore be reduced to an absolute minimum. Since the aviator does not want to write, it is necessary that he should simply note the terms to be added upon a graduated dial which permits him to read off the total.

The aim of the aeronautical ephemerides is to reduce to a minimum the determination of the astronomical data required in subsequent calculations for fixing the position.

Graduated dials have been constructed for facilitating the additions, but not for subtractions. The latter cannot be easily effected by the aviator and should not be required of him. Thus, in calculating the position by the astronomical method, he writes as follows :—

$$\text{Hour Angle} = \text{Sidereal Time} - \text{Right Ascension.}$$

a formula which requires subtraction. Therefore it is necessary to furnish him with tables giving the verse-ascension, which is the right ascension subtracted from 24 hours — this in degrees and minutes of degrees, since

1887. -- Lundi, AVRIL 19.

Table with columns: TEMPS de Greenwich, VENUS, MARS, JUPITER, SATURNE, TEMPS de Greenwich, VENUS, JUPITER, SATURNE. Includes sub-columns for Ciel, Hauteur, Décl., Anom., and Dist. from Earth.

1887. -- Lundi, AVRIL 19.

Table with columns: TEMPS de Greenwich, SOLEIL, LUNE, TEMPS de Greenwich, SOLEIL, LUNE. Includes sub-columns for Ciel, Hauteur, Décl., Anom., and Dist. from Earth.

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the sidereal time is measured in degrees and minutes of arc. As the aviator is no longer using units graduated in time intervals it is necessary to furnish him with these data in degrees and minutes of arc.

The aeronautical ephemerides have been compiled with the sole aim of satisfying these requirements of aviation. We desire that the aviator should be able to obtain the data required for calculating his position in the simplest manner possible. After that he will complete his calculations with the aid of other devices, but these tables have been compiled so that by a simple addition one may obtain the hour angle immediately.

There exist also other precedents. In America, in the United States, there were published in 1932 ephemerides of this kind which were not entirely identical but which were nevertheless compiled especially for aviators.

In France, at the close of the year 1935 the Air Ministry requested the *Bureau des Longitudes* to publish tables of this kind and I was given charge of their preparation with the assistance of the expert calculators in this Bureau; it sufficed for me to state what we wanted to accomplish and they produced something with which I could not find the slightest fault. With the exception of the civil time, which is given in hours, all of the data for the calculations are given in degrees and minutes of arc. Since the aviators wish to reduce all calculations to a minimum, it is necessary to avoid interpolations. With the aviation sextants the closest reading at present is to within about 6 or 7 minutes of arc. It will suffice therefore to give the data to the nearest minute of arc and, for the Sun, for intervals of 5° in order that no interpolations should be necessary. The aviator has only to take from the table the number corresponding to the time of the observation in order to calculate the hour angle. This is obtained accurately in the case of the Sun, Venus, Mars, Jupiter and Saturn : but for the Moon, which aviators are led to observe rather frequently, we have not been able to do away entirely with interpolations. For the Moon, therefore, the problem still remains.

For practical reasons, one should not require aviators to carry such a large volume with them on their flights, and therefore we have perforated the pages so that those actually needed for the data may be torn out of the volume. If, for instance, one were leaving on 14th August for a flight of two days, one would simply take the pages for August 14th and 15th, together with the appended fascicle which comprises only 5 pages.

This appendix contains the positions of 72 principal stars given for the middle of the year, and a few small secondary tables as well. A single page gives the calculations for latitude from the observed altitude of Polaris. There is also a correction for refraction. At an altitude of 10,000 metres, the refraction is only about three-tenths of the value on the ground.

The appendix also gives the dip of the horizon, which is useful at times; the semi-diameter of the Sun and the semi-diameter of the Moon. We have been obliged to insert a table giving the parallax of the Moon.

The first volume was issued in 1936, and the publication will be regularly continued. In the ephemerides of 1937 certain simplifications have been effected. For the parallax of the Moon another table has been provided in order to avoid the necessity for multiplying by the sine of the altitude.

These tables have been published at the request of the Air Ministry; we hope that civil aviation will also avail itself of these tables; the purchase of them in this connection has already begun.

I might add that it is preferable to make use of chronometers graduated in degrees and minutes. In order that these instruments may be easily read in spite of the agitations, one may employ the type known as the "guichet" (window or gas-meter type) in which the degrees appear in openings, or in large figures.

RECENT CHARTS.

Subsequent to the International Hydrographic Conference of 1932, the French *Service Hydrographique* adopted for its charts, on which the metre is used as unit, the system of representation of the isobaths recommended by that Conference. A certain number of charts have been prepared in accordance with this convention. Here is one that has just been completed (Chart N° 5818). Here is another now being engraved. The convention is readily understood by reference to the table on page 410 of the Report of Proceedings of the 1932 Hydrographic Conference.

